



IEUVI Mask TWG Panel

Buried defects investigated with simulation

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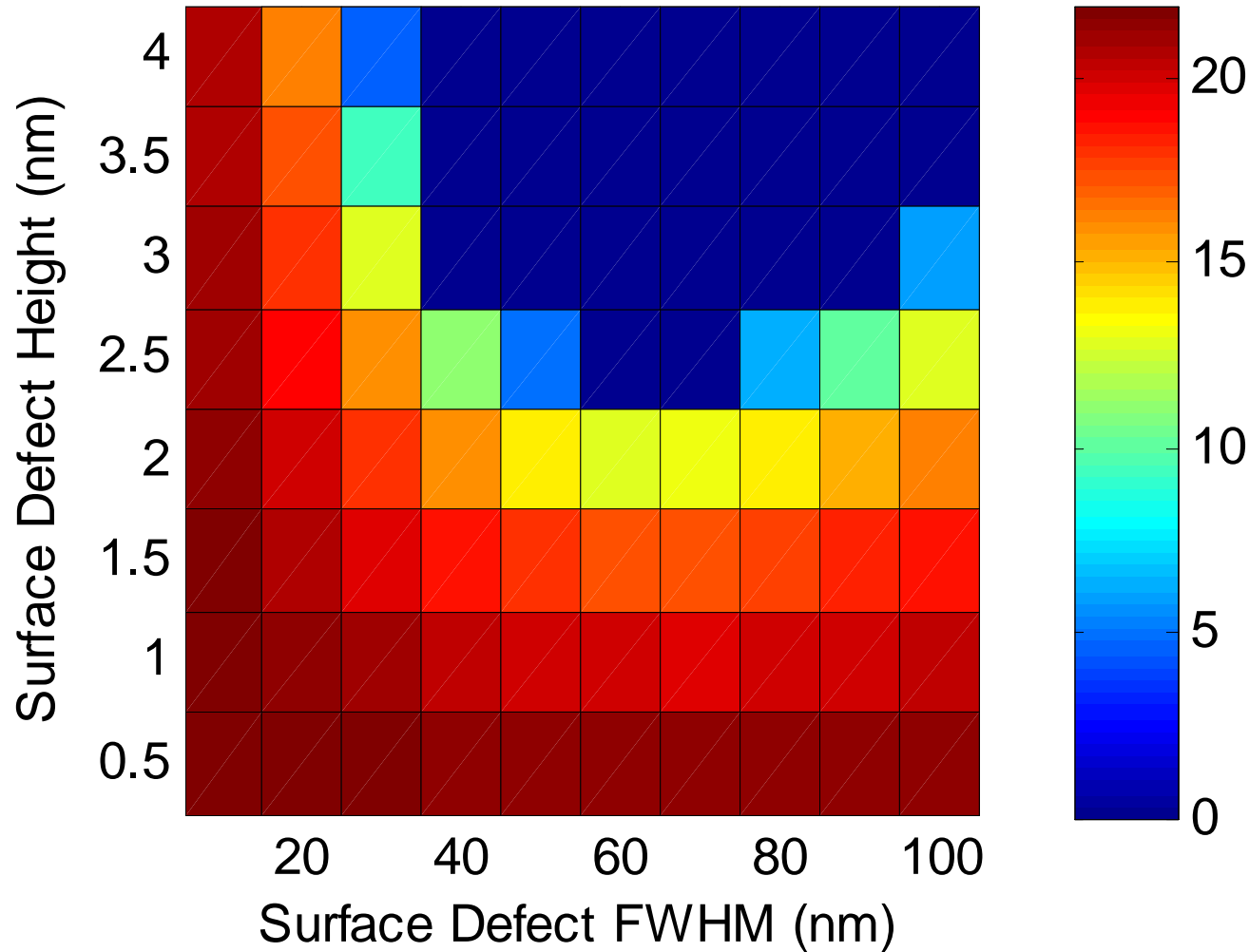
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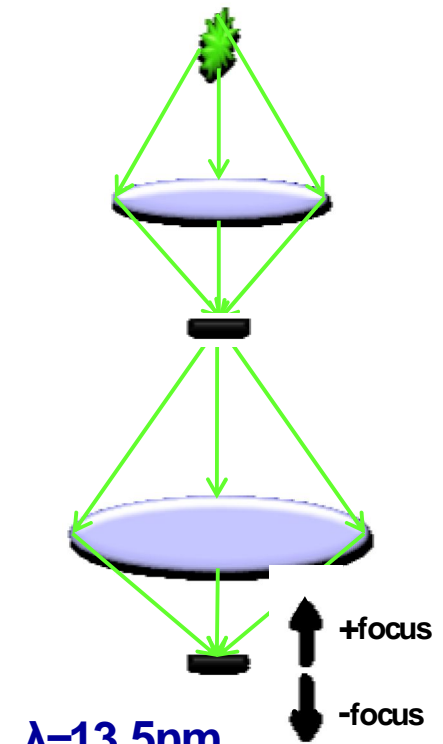


65nm is the worst case surface FWHM

CD Change for 22nm lines with defect located 25nm from the shadowed absorber edge



Optical Model



$\lambda=13.5\text{nm}$

$\text{NA}=0.32$

4x Reduction

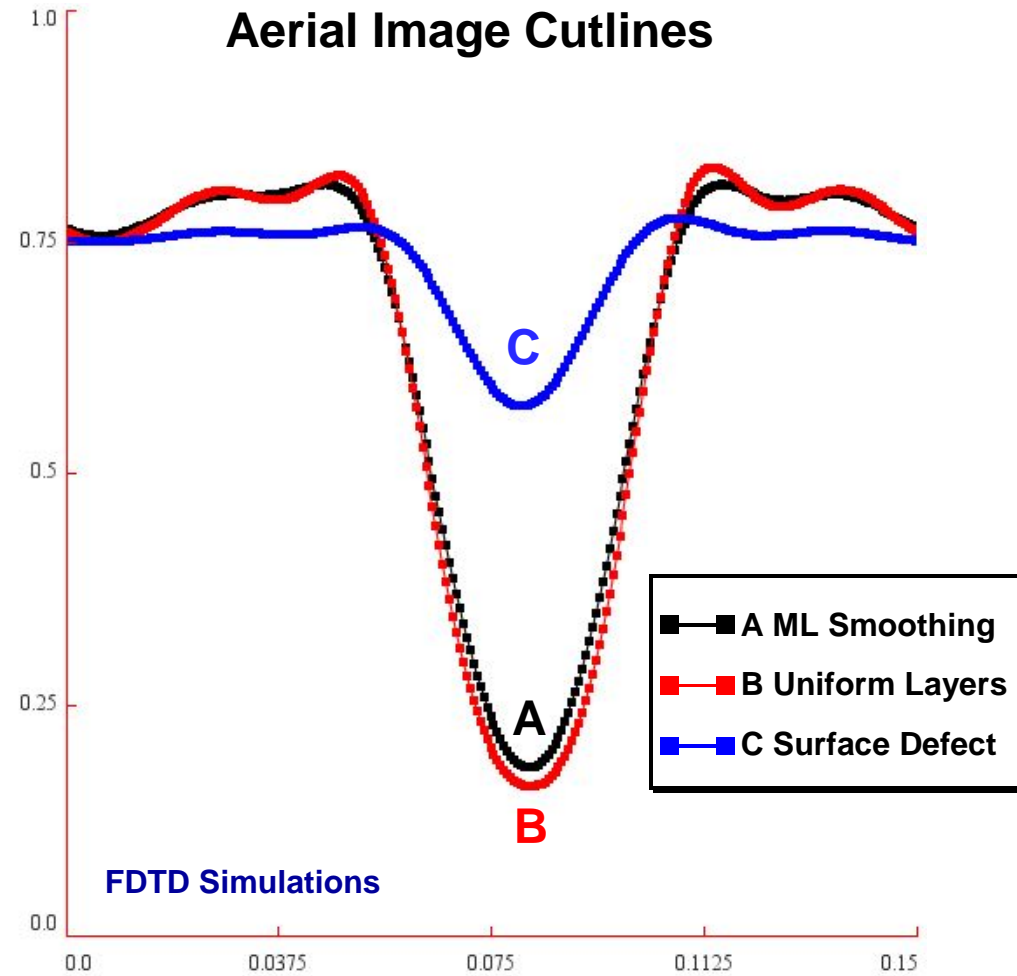
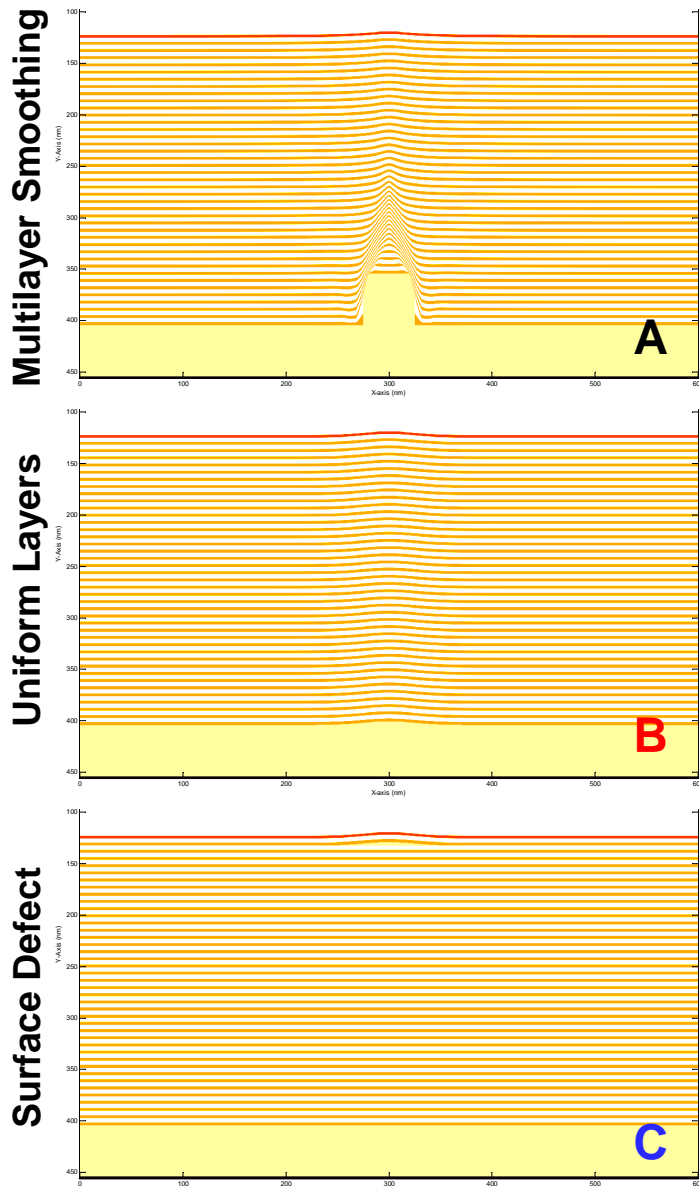
$\sigma=0.75$ Except for final section on illumination

6° incident light



What is below the multilayer surface?

Three example geometries, all with a surface height of 3.55nm, and FWHM of 59nm ($3.55\text{nm} = 189^\circ$)

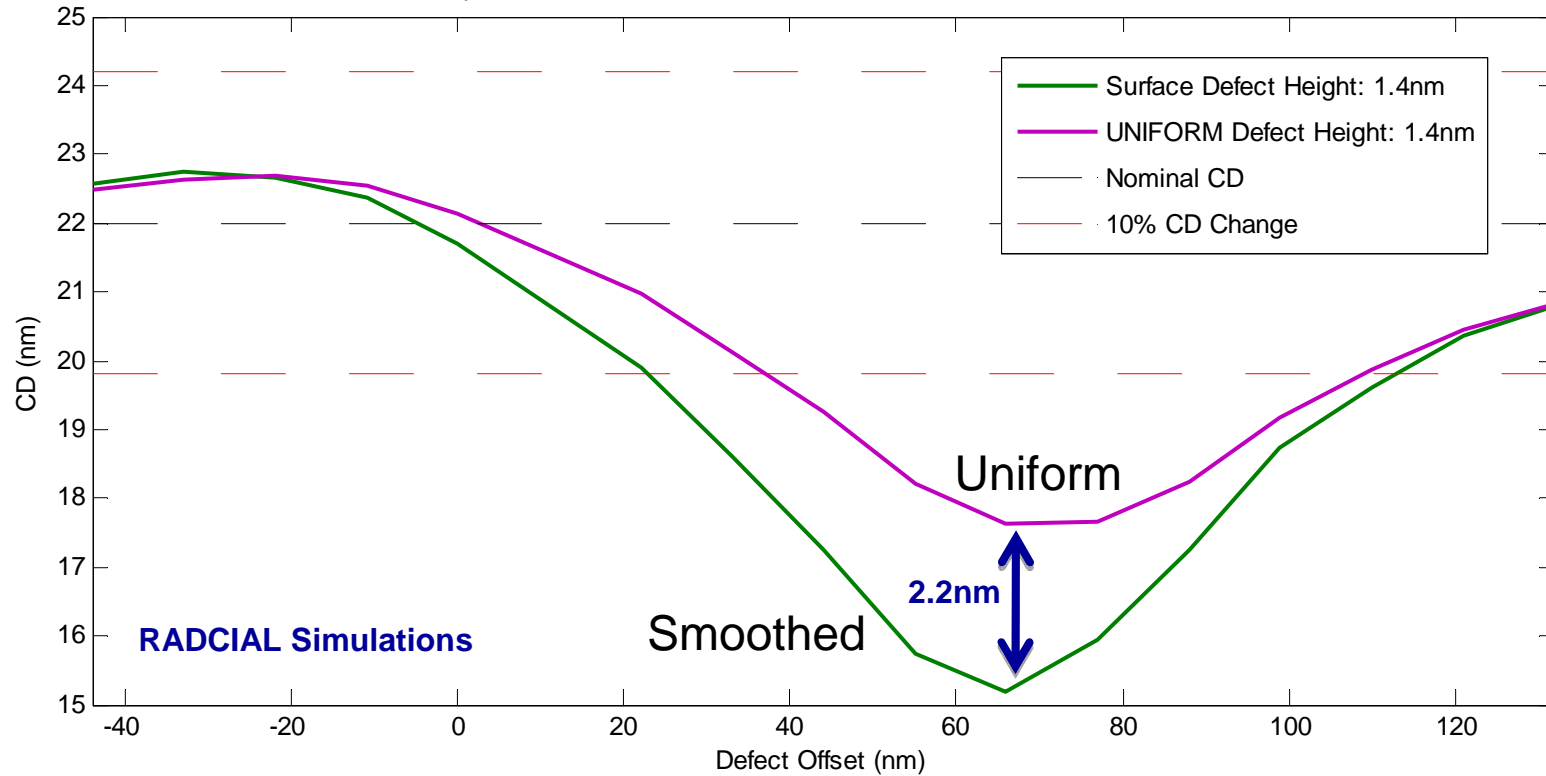


For the rest of the presentation, all multilayers are smoothed like the TOP figure unless noted otherwise



Multilayer has a larger effect for defects near features

Space CD as a Function of Defect Position for 22nm Dense Lines

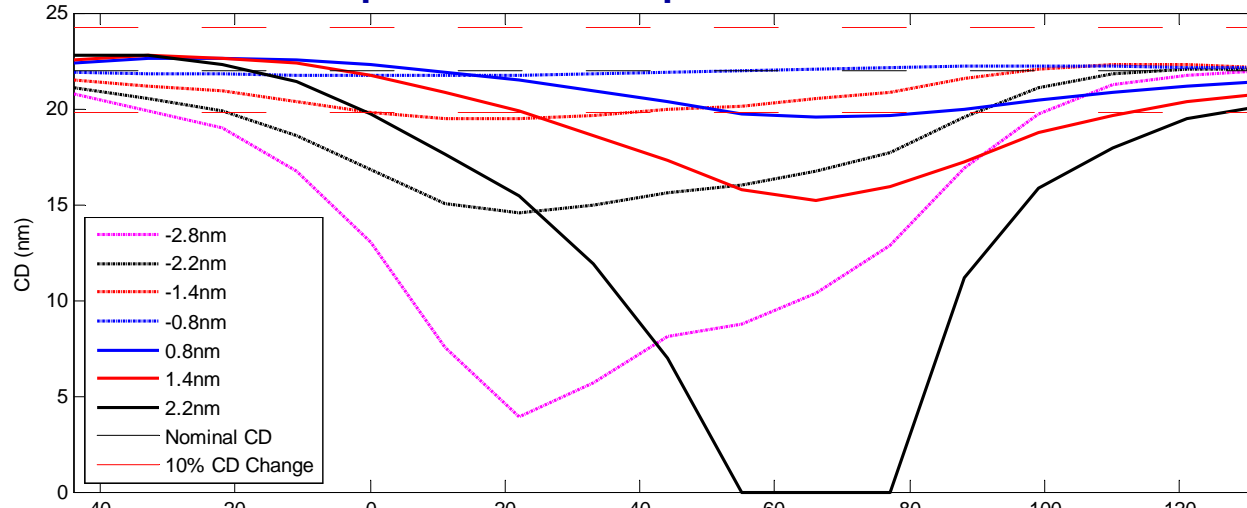


	Surface Height	Substrate Height
Smoothed	1.4nm	5nm
Uniform	1.4nm	1.4nm



Summary of Buried Defect Printability for 22nm Dense Lines

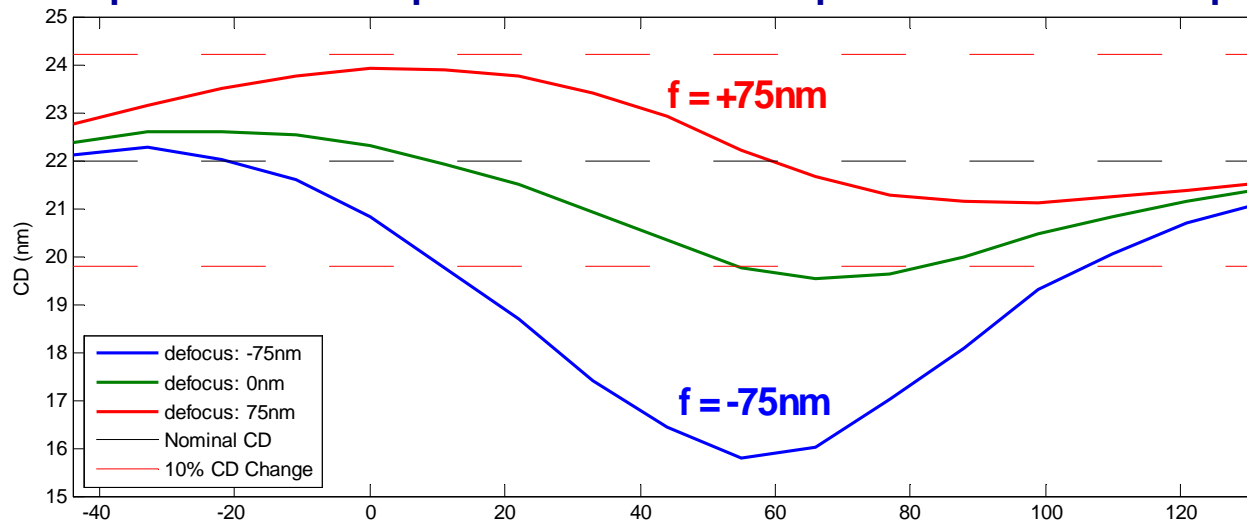
Space CD vs. defect position at best focus



- Maximum allowable surface bump: 0.8nm
- Maximum allowable surface pit: 1.4nm
 - No ML smoothing for pits

- Pits have opposite worst case position
- I believe worst case position is determined by the relative phase related to nonzero k_x caused by 6° incident light

Space CD vs. defect position at various focus positions for 0.8nm bump



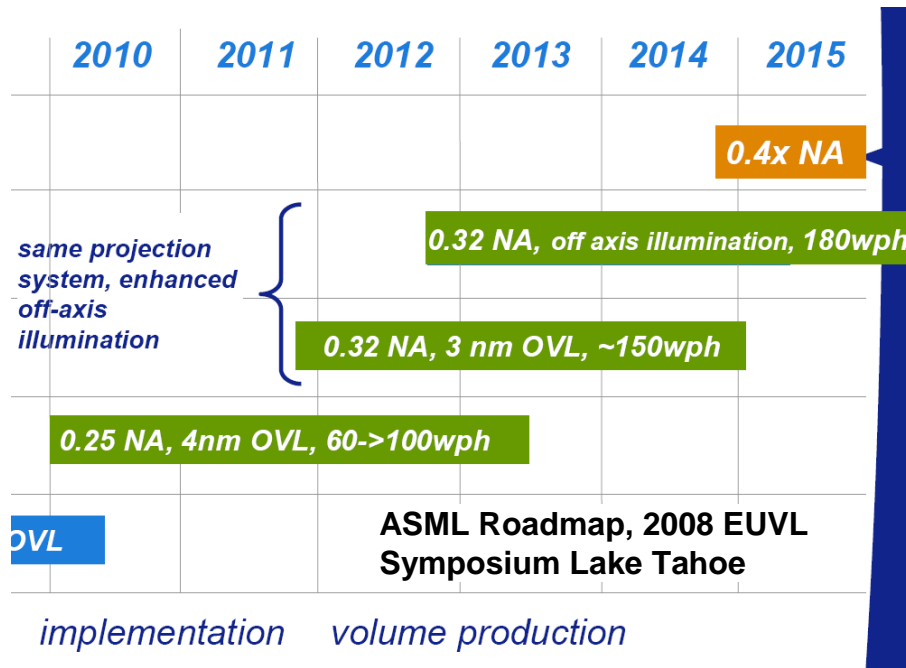
- Allowable defect size depends on expected focus variation
- Already verified by experiments

Defect under absorber Defect under absorber edge Defect in center of space Defect under absorber edge Defect under absorber

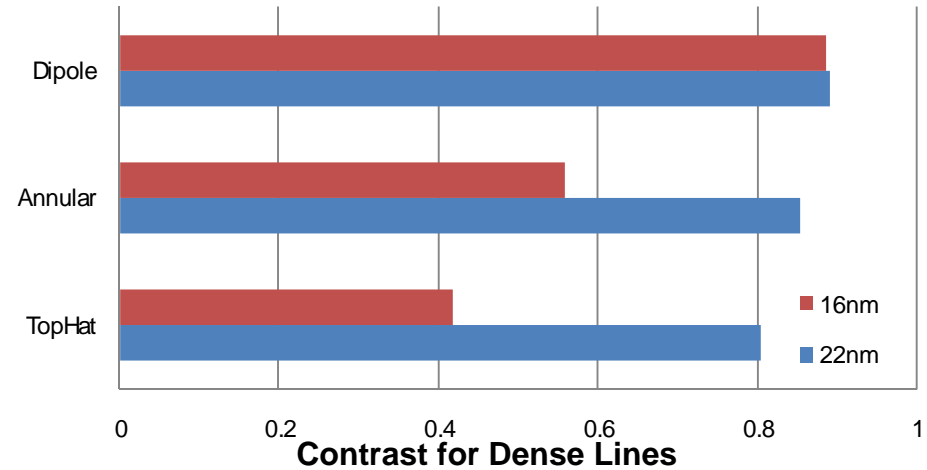
RADCIAL Simulations



Illumination Effects Defect Printability



Off-axis illumination will likely be used when HVM EUVL is introduced



Images of an isolated defects

